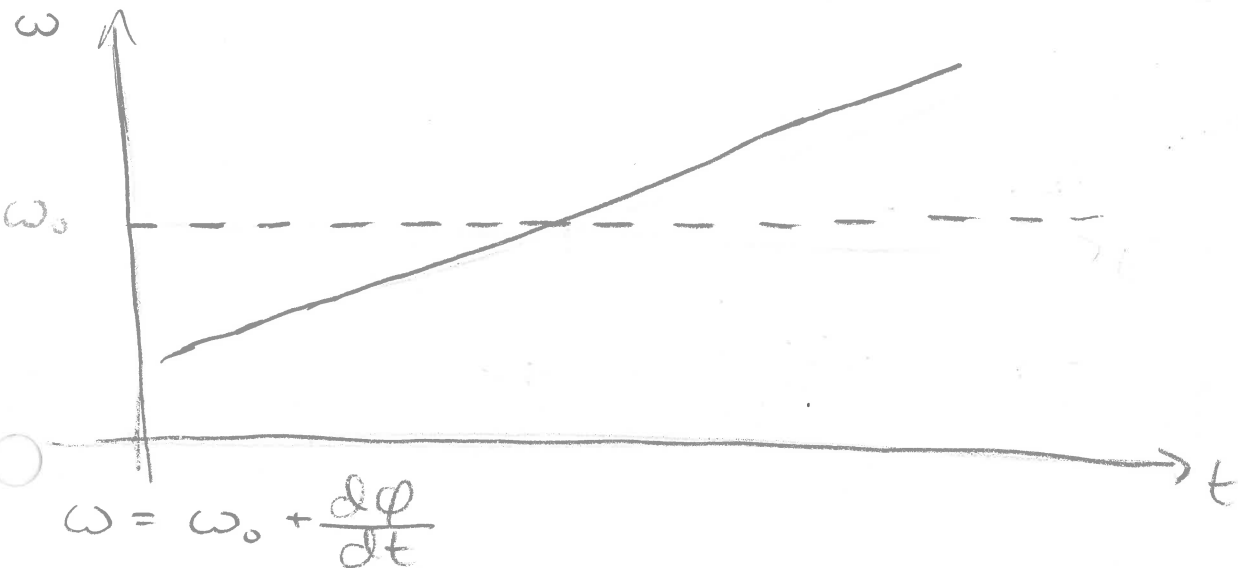
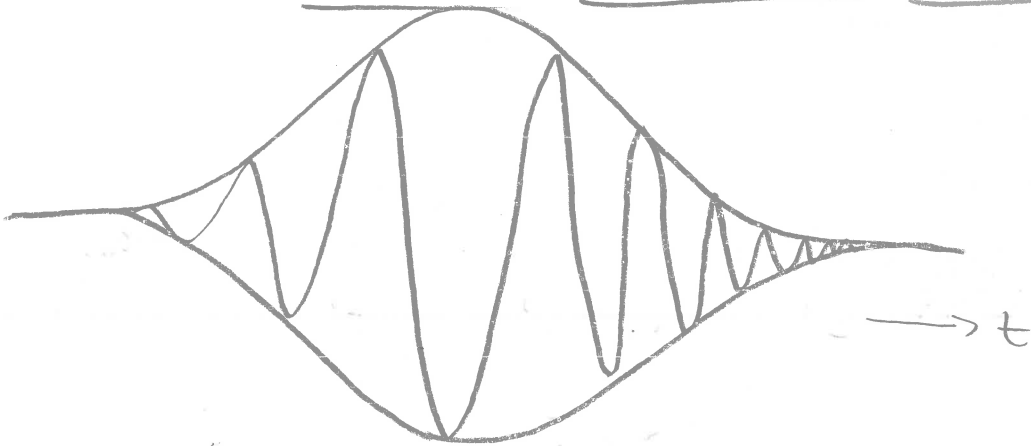


Final Exam Solutions

1.
(a.)



$$\frac{d\omega}{dt} = \frac{d^2\varphi}{dt^2} > 0 \rightarrow \text{up-chirped pulse}$$

(b.)

$$\vec{E} = \frac{1}{2} E_0 e^{-\frac{x^2+y^2}{w^2(z)}} e^{-(1+ia)\frac{t^2}{\tau_0^2}} e^{i(\omega_0 t - kz)} e^{i\varphi_0} \hat{x} + \text{c.c.}$$

(c.)

$$I(t) = \frac{1}{2} \epsilon_0 c n \vec{E}^* \vec{E} = \frac{\epsilon_0 c}{2} E^* E$$

$$E^* E = \left(\frac{1}{2}\right)^2 E_0^2 e^{-\frac{2(x^2+y^2)}{w^2(z)}} \left[e^{-(1-ia)\frac{t^2}{\tau_0^2}} e^{-i(\omega_0 t - kz)} e^{-i\varphi_0} e^{-(1+ia)\frac{t^2}{\tau_0^2}} e^{i(\omega_0 t - kz)} e^{i\varphi_0} \right]$$

$$E^*E = \frac{E_0^2}{4} e^{-\frac{2(x^2+y^2)}{w^2(z)}} \left[2e^{-\frac{2t^2}{\tau_0^2}} + e^{-2(1-ia)\frac{t^2}{\tau_0^2}} e^{-2i(\omega_0 t - kz) - 2i\phi_0} + e^{-2(1+ia)\frac{t^2}{\tau_0^2}} e^{2i(\omega_0 t - kz) + 2i\phi_0} \right]$$

$$= \frac{E_0^2}{4} e^{-\frac{2(x^2+y^2)}{w^2(z)}} e^{-\frac{2t^2}{\tau_0^2}} \left[2 + e^{+2ia\frac{t^2}{\tau_0^2} + 2i(kz - \omega_0 t) + \phi_0} + e^{-2ia\frac{t^2}{\tau_0^2} - 2i(kz - \omega_0 t) - \phi_0} \right]$$

$$= \frac{E_0^2}{4} e^{-\frac{2(x^2+y^2)}{w^2(z)}} e^{-\frac{2t^2}{\tau_0^2}} \left\{ 2 \left[1 + \cos\left(\frac{2at^2}{\tau_0^2} + kz - \omega_0 t - \phi_0\right) \right] \right\}$$

$$E^*E = \frac{E_0^2}{2} e^{-\frac{2(x^2+y^2)}{w^2(z)}} e^{-\frac{2t^2}{\tau_0^2}} \left[1 + \cos\left\{ -\left(\omega_0 - \frac{2at}{\tau_0^2}\right)t - kz + \phi_0 \right\} \right]$$

$$I(t) = \frac{\epsilon_0 c E_0^2}{4} e^{-\frac{2(x^2+y^2)}{w^2(z)}} e^{-\frac{2t^2}{\tau_0^2}} \left[1 + \cos\left[\underbrace{\left(\omega_0 - \frac{2at}{\tau_0^2}\right)t}_{\omega(t)} - kz + \phi_0 \right] \right]$$

$\cos(-\theta) = \cos\theta$

$$\omega = \omega_0 - \frac{2at}{\tau_0^2}$$

Depending on the sign of a , $\omega(t)$ either increases or decreases w/ t . For our up-chirped pulse, we know that $\omega(t)$ increases w/ t , so a is negative.

2. (a.)

$$I_p = \frac{2W}{10^3 \text{ Hz}} \frac{1}{\pi (10^{-5} \text{ m})^2} \frac{1}{\tau_G} = \frac{6.37 \times 10^6}{\tau_G} \frac{\text{W}}{\text{m}^2}$$

$$I_p = \frac{cn\epsilon_0}{2} E_p^2, \text{ where } E_p = 5 \times 10^{11} \text{ V/m}, n=1$$

$$\frac{6.37 \times 10^6 \frac{\text{W}}{\text{m}^2}}{\tau_G} = \frac{3 \times 10^8 \text{ m/s} \times 8.85 \times 10^{-12} \text{ F/m}}{2} 25 \times 10^{22} \text{ V}^2/\text{m}^2$$

$$\tau_G = 19.2 \text{ fs}$$

(b.)

$$\Delta\nu_G \tau_G = 0.441 \quad (\text{I'll also accept } \Delta\nu_G \tau_G = \frac{1}{2})$$

$$\Delta\nu_G = \frac{0.441}{\tau_G}$$

$$\Delta\nu_G \tau_G = 1$$

$$\frac{c}{\lambda} = \nu \rightarrow \frac{c}{\lambda^2} \Delta\lambda = \Delta\nu$$

$$\Delta\lambda = \frac{\lambda^2}{c} \Delta\nu$$

$$\Delta\lambda = \frac{\lambda^2}{c} \frac{0.441}{\tau_G} = \frac{(8 \times 10^{-7} \text{ m})^2}{3 \times 10^8 \text{ m/s}} \frac{0.441}{1.92 \times 10^{-14} \text{ s}}$$

$$\Delta\lambda = 4.90 \times 10^{-8} \text{ m} = 49 \text{ nm}$$

$$\lambda = 111 \text{ nm}, \Delta\nu_G \tau_G = 1$$

$$\lambda = 55.6 \text{ nm}, \Delta\nu_G \tau_G = \frac{1}{2}$$

(c.)

$$\frac{2W}{10^3 \text{ Hz}} = 2 \times 10^{-3} \text{ J/pulse} = 2 \times 10^{-3} \text{ J/pulse} \times \frac{1 \text{ eV}}{1.603 \times 10^{-19} \text{ J}} = 1.25 \times 10^{16} \frac{\text{eV}}{\text{pulse}}$$

$$\frac{hc}{\lambda_0} = \frac{1240 \text{ nm-eV}}{800 \text{ nm}} = 1.55 \text{ eV/photon (on average)}$$

$$\frac{1.25 \times 10^{16} \text{ eV/pulse}}{1.55 \text{ eV/photon}} = 8.05 \times 10^{15} \text{ photons/pulse}$$

3.

(a) $\Delta \nu_c = \frac{FSR}{\mathcal{F}}$

$$FSR = \frac{c}{2nL} = \frac{3 \times 10^8 \text{ m/s}}{2 \times 1.25 \times 0.40 \text{ m}} = 3 \times 10^8 \text{ Hz}$$

$$\mathcal{F} = \frac{\pi \sqrt{R}}{1-R} = \frac{\pi \sqrt{0.99}}{0.01} = 3.126 \times 10^2$$

$$\Delta \nu_c = \frac{FSR}{\mathcal{F}} = \frac{1.5 \times 10^8 \text{ Hz}}{3.126 \times 10^2} = \boxed{9.60 \times 10^5 \text{ Hz}}$$

(b) $\Delta \nu_L = \frac{(2)\pi h \nu_L \Delta \nu_c^2}{P_L} = \frac{(2)\pi (6.626 \times 10^{-34} \text{ J-s}) (4.74 \times 10^{14} \text{ Hz}) (9.60 \times 10^5 \text{ Hz})^2}{10^{-2} \text{ W}}$

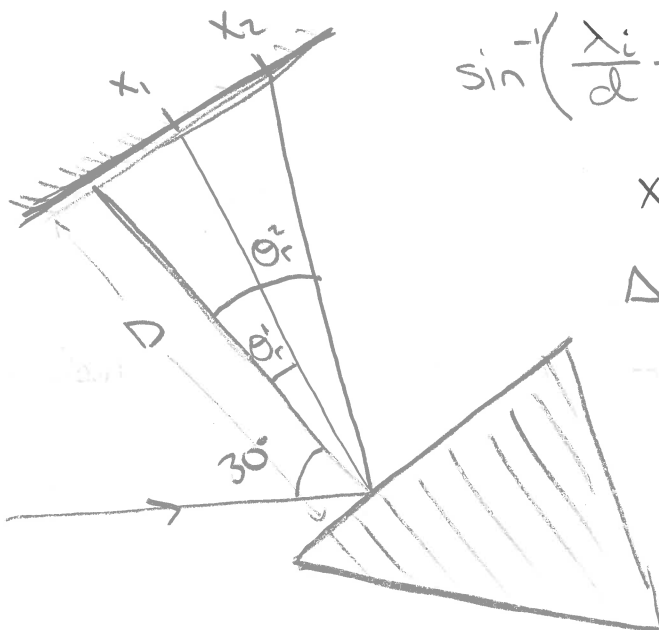
$$\left[\nu_L = \frac{c}{\lambda} = \frac{3 \times 10^8 \text{ m/s}}{6.328 \times 10^{-7} \text{ m}} = 4.74 \times 10^{14} \text{ Hz} \right]$$

$$\Delta \nu_L = 1.82 \times 10^{-4} \text{ Hz} = \boxed{182 \mu\text{Hz}}$$

$$\Delta \nu_L = 9.09 \times 10^{-5} \text{ Hz} = \boxed{90.9 \mu\text{Hz}}$$

if you used a factor of 2
if you used a factor of 1

(c.)



$$\sin^{-1}\left(\frac{x_i}{d} - \frac{1}{2}\right) = \theta_r^i$$

$$x_i = D \tan \theta_r^i$$

$$\Delta x = x_2 - x_1 = D(\tan \theta_r^2 - \tan \theta_r^1)$$

Since $\lambda_0 = 632.8 \text{ nm} \Rightarrow \nu_L = 4.74 \times 10^{14} \text{ Hz}$

and BW of $\Delta\nu_L = 181 \text{ MHz}$ is approximately 18 orders of magnitude smaller, we are going to be dealing with a hard to measure linewidth.

$$\Delta x = 20 \mu\text{m} / 2 = 10^{-5} \text{ m}$$

$$D = 500 \times 10^{-3} \text{ m} = 0.5 \text{ m}$$

$$d = \frac{10^{-3}}{2000} = 5 \times 10^{-7} \text{ m}$$

$$\lambda_1 = 632.8 \text{ nm}$$

$$\theta'_r = \sin^{-1} \left(\frac{\lambda_1}{d} - \frac{1}{2} \right)$$

$$\theta'_r = 0.872$$

$$\Delta x = D (\tan \theta_r^2 - \tan \theta'_r)$$

$$\sin^{-1} \left(\frac{\Delta x}{D} + \tan \theta'_r \right) = \theta_r^2$$

$$\text{But } \theta_r^2 = \sin^{-1} \left(\frac{\lambda_2}{d} - \frac{1}{2} \right)$$

$$\sin^{-1} \left(\frac{\lambda_2}{d} - \frac{1}{2} \right) = \tan^{-1} \left(\frac{\Delta x}{D} + \tan \theta'_r \right)$$

$$\lambda_2 = d \left[\sin \left(\tan^{-1} \left(\frac{\Delta x}{D} + \tan \theta'_r \right) \right) + \frac{1}{2} \right]$$

$$= 5 \times 10^{-7} \text{ m} \left[\sin \left(\tan^{-1} \left[\frac{10^{-5} \text{ m}}{0.15 \text{ m}} + \tan(0.872) \right] \right) + \frac{1}{2} \right]$$

$$\lambda_2 = 632.82 \text{ nm}$$

If $\Delta\lambda > 2(632.82 \text{ nm} - 632.8 \text{ nm})$ or $\Delta\lambda > 0.04 \text{ nm}$, then

we can resolve it on more than one pixel.

However, $\Delta\nu_L = 181 \text{ MHz}$ gives us a bandwidth of $\Delta\lambda_L = \frac{c}{\nu_L^2} \Delta\nu_L$ or $2.4 \times 10^{-25} \text{ m}$, which is beyond any known measurement technique.

4.

(a.) $\tau_0 = 5 \text{ fs}$, $\lambda_0 = 500 \text{ nm}$, $n = 1$

Method 1

$$c \Delta t = \Delta x$$

$$3 \times 10^8 \text{ m/s} \times 5 \times 10^{-15} \text{ s} = \Delta x$$

$$3 \times 5 \times 10^{-7} \text{ m} = \Delta x$$

Method 2

$$\text{Period} = \frac{1}{\nu} = \frac{\lambda}{c} = \frac{5 \times 10^{-7} \text{ m}}{3 \times 10^8 \text{ m/s}} = 1.67 \times 10^{-15} \text{ s} = 1.67 \text{ fs}$$

$$\text{Pulsewidth} = 5 \text{ fs}$$

$$\frac{\text{Pulsewidth}}{\text{Period}} = \frac{5 \text{ fs}}{1.67 \text{ fs}} = 3.0 \text{ optical cycles}$$

$$\frac{\Delta x}{\lambda} = \# \text{ of optical cycles in one pulsewidth}$$

$$= \frac{3 \times 5 \times 10^{-7} \text{ m}}{5 \times 10^{-7} \text{ m}} = 3.0 \text{ optical cycles}$$

← standard definition

$$\frac{2\Delta x}{\lambda} = 6.0 \text{ optical cycles in two pulsewidths (also acceptable)}$$

(b.)

For every pulsewidth, τ_0 , the carrier is oscillating

just 1.5 optical cycles. The temporal portion of our envelope, $e^{-\frac{t^2}{\tau_0^2}}$, changes by e^{-1} (factor of 2.72) in magnitude for every 1.5 carrier oscillations.

SVEA (in temporal form):

$$\left| \frac{d\tilde{E}(t)}{dt} \right| \ll \omega_0 |\tilde{E}(t)|$$

$$\left| \frac{-t}{\tau_0^2} e^{-\frac{t^2}{\tau_0^2}} \right| \ll \frac{2\pi c}{\lambda_0} \left| e^{-\frac{t^2}{\tau_0^2}} \right|$$

At $t = \tau_0$

$$\frac{1}{\tau_0} \ll \frac{2\pi \cdot 3 \times 10^8 \text{ m/s}}{5 \times 10^{-7} \text{ m}}$$

$$\frac{1}{5 \times 10^{-15} \text{ Hz}} \ll \pi \cdot 10^{15} \text{ Hz} \cdot \frac{6}{5}$$

$$1 \ll 6\pi \rightarrow 1 \ll 18$$

Although it is close, we are still in the limit that the envelope is slowly varying with respect to the carrier wave.

